

Déformations en relativité poincaréenne et galiléenne : les tenseurs de conformation et de friction

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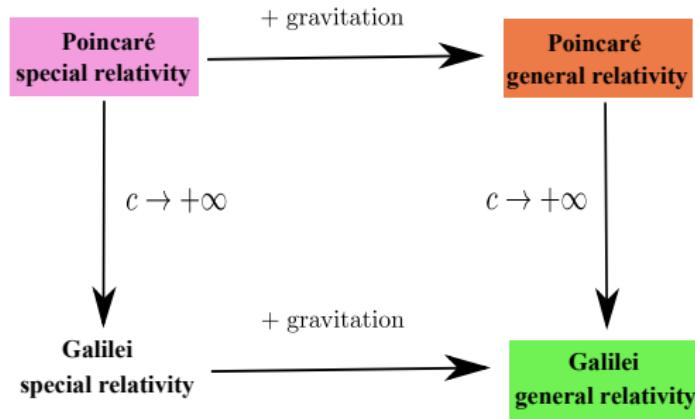


Two deformation tensors in Relativity

- Relativistic reversible media Jean-Marie Souriau
Géométrie et Relativité (1964)
Conformation tensor
- Relativistic dissipative media Jean-Marie Souriau
Lect. Notes in Math. 676 (1976)
Friction tensor

Galilei and Poincaré relativities

Christian Cardall terminology :
A unified perspective on
Poincaré and Galilei relativity 2023



Galilei group and geometry

Galilei general relativity

- **Chart** : $\mathbb{R}^4 \ni X = \begin{pmatrix} t \\ x \end{pmatrix} \mapsto \mathbf{X} \in \mathcal{M}$ space-time
- **4-velocity** : \vec{U} represented by $U = \frac{dX}{dt} = \begin{pmatrix} 1 \\ v \end{pmatrix}$
- **Galilei group** GAL is the one of the **Galilean transformations** :

$$X = P X' + C \quad \text{with} \quad P = \begin{pmatrix} 1 & 0 \\ \mathbf{u} & R \end{pmatrix}, \quad C = \begin{pmatrix} \tau_0 \\ k \end{pmatrix}$$

where $\mathbf{u} \in \mathbb{R}^3$ is the **Galilean boost** and R is a rotation

- **Clock form** $\tau = dt$ represented by the Galilean invariant $(1 \ 0^T)$

Galilei group and geometry

- The closure condition $d\omega = 0$ of the presymplectic form gives

$$\operatorname{curl} g + 2 \frac{\partial \Omega}{\partial t} = 0, \quad \operatorname{div} \Omega = 0$$

then there exist **potentials** ϕ, A generating

the **gravity** $g = -\operatorname{grad} \phi - \frac{\partial A}{\partial t}$ and **Coriolis' effect** $\Omega = \frac{1}{2} \operatorname{curl} A$

- The corresponding Lagrangian

$$\mathcal{L} = \frac{1}{2} m \| v \|^2 + m A \cdot v - m \phi$$

is covariant provided

$$\phi' = \phi - A \cdot u - \frac{1}{2} \| u \|^2, \quad A' = R^T(A + u)$$

Spacetime metrics

Poincaré general relativity

For weak gravitational fields

- $G = \begin{pmatrix} c^2 + 2\phi & -A^T \\ -A & -1_{\mathbb{R}^3} \end{pmatrix} = \epsilon^{-1} \begin{pmatrix} (-1) \\ G^{(0)} \end{pmatrix}$

where $\epsilon = c^{-2}$, $G^{(-1)} = dt \otimes dt$ and $G^{(0)}$ are Galilean tensors

- Truncated expansion of

$$G^{-1} \cong \begin{pmatrix} 0 & 0 \\ 0 & -1_{\mathbb{R}^3} \end{pmatrix} + \epsilon \begin{pmatrix} 1 & -A^T \\ -A & AA^T \end{pmatrix}$$

Each term is a Galilean 2-contravariant tensor

- $\sqrt{-\det G} \cong c \left(1 + \epsilon \left(\phi + \frac{\|A\|^2}{2} \right) \right)$
Galilean invariant

Hilbert-Einstein functional $\mathcal{A} = \mathcal{A}_M + \mathcal{A}_G = \int (p_M + p_G) \sqrt{-\det G} d^4X$

Adjoint map

Let $(\mathcal{M}_0, \mathbf{G}_0), (\mathcal{M}, \mathbf{G})$ be two Riemannian spaces

The **adjoint** of a map

$$\mathbf{A} : T_{x_0} \mathcal{M}_0 \rightarrow T_x \mathcal{M}$$

is the map

$$\mathbf{A}^* : T_x \mathcal{M} \rightarrow T_{x_0} \mathcal{M}_0$$

such that

$$\mathbf{G}(\mathbf{A}(\vec{U}_0), \vec{V}) = \mathbf{G}_0(\vec{U}_0, \mathbf{A}^*(\vec{V}))$$

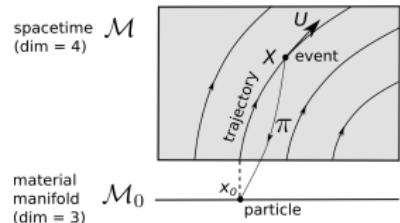
In a local chart, if \mathbf{A} and \mathbf{A}^* are represented by the matrices A and A^*

$$A^* = G^{-1} A^T G_0$$

where G_0 and G are Gram's matrices of \mathbf{G}_0 and \mathbf{G}

Reversible material : conformation tensor

line bundle : $\pi : \mathcal{M} \rightarrow \mathcal{M}_0$
 represented by : $x_0 = \pi(X)$



- **Conformation tensor** [Souriau 1964 Géométrie & Relativité]

$$\mathcal{D} = -\frac{\partial x_0}{\partial X} \left(\frac{\partial x_0}{\partial X} \right)^* = -\frac{\partial x_0}{\partial X} G^{-1} \left(\frac{\partial x_0}{\partial X} \right)^T G_0$$

where $\frac{\partial x_0}{\partial X} = (-F^{-1} v, F^{-1})$ with $F = \frac{\partial x}{\partial x_0}$

- For an **isotropic material**, $G_0 = 1_{\mathbb{R}^3}$
- mass density (Galilean invariant) $\rho = \frac{\rho_0(x_0)}{\det F}$
- $\mathcal{D} \cong F^{-1} (1_{\mathbb{R}^3} - \epsilon(v + A)(v + A)^T) F^{-T}$

Conformation tensor

Galilei general relativity

Hypothesis (Souriau) : the functional due to the matter is a function of the conformation tensor.

For the simplest matter (dust), we claim that

$$\rho_M \sqrt{-\det G} = c^2 \rho_0(x_0) \sqrt{\det \mathcal{D}(-\det G)} \cong c [\rho c^2 - \rho \left(\frac{\|v\|^2}{2} + A \cdot v - \phi \right)]$$

energy	Lagrangian
at rest	density

In Relativity, the three following phenomena

- the deformation (in the simplest form, contained in the density)
- the kinetic
- and the gravitation

are coupled through a unique tensor, the **conformation**

Conformation tensor

Poincaré special relativity

Comment : if $F = \mathbf{1}_{\mathbb{R}^3}$ and $\rho_0 = C^{te}$,
we recover the proper-time (exactly)

$$\mathcal{A}_M = c \cdot m_0 c^2 \int \sqrt{1 - \|v\|^2/c^2} dt$$

Galilei general relativity

More general case

$$\rho_M \sqrt{-\det G} = \left[c^2 \rho_0(x_0) \sqrt{\det \mathcal{D}} + \mathcal{W}(x_0, \mathcal{D}) \right] \sqrt{-\det G}$$

Neo-Hookean hyperelastic material $\mathcal{W}(x_0, \mathcal{D}) = \mu \operatorname{Tr}(\mathcal{D}^{-1})$

$$\begin{aligned} \rho_M \sqrt{-\det G} &\cong c [\rho \mathbf{c}^2 - \rho (\frac{\|v\|^2}{2} + A \cdot v - \phi)] + \mu \operatorname{Tr}(C) \\ &\quad + \frac{\mu}{c^2} \left\{ \frac{\|F^T v\|^2}{1 - \|v\|^2/c^2} + \operatorname{Tr}(C) \left(\phi + \frac{\|A\|^2}{2} \right) \right\} \end{aligned}$$

with $C = F^T F$

Dissipative material : friction tensor

Poincaré general relativity

- Relativistic Thermodynamics Jean-Marie Souriau
Lect. Notes in Math. 676 (1976)
- Dissipative constitutive laws in Special Relativity
Claude Vallée, IJES (1981)

Galilei general relativity

- Book with Claude Vallée :
Galilean Mechanics and Thermodynamics of Continua
(ISTE-Wiley, 2016)

Bargmannian transformations

Galilei general relativity

- The space-time \mathcal{M} is embedded into a space $\hat{\mathcal{M}}$ of dimension 5 :
 $\mathcal{M} \rightarrow \hat{\mathcal{M}} : \mathbf{X} \mapsto \hat{\mathbf{X}} = \hat{f}(\mathbf{X})$
- We built a group of affine transformations $\hat{\mathbf{X}}' \mapsto \hat{\mathbf{X}} = \hat{P}\hat{\mathbf{X}}' + \hat{C}$ of \mathbb{R}^5 which are Galilean when acting onto the space-time of which the linear part is :

$$\hat{P} = \begin{pmatrix} 1 & 0 & 0 \\ u & R & 0 \\ \frac{1}{2} \|u\|^2 & u^T R & 1 \end{pmatrix}$$

Their set is the **Bargmann's group**,

Temperature 5-vector

- The reciprocal temperature $\beta = 1/\theta$ is generalized as a Bargmannian 5-vector :

$$\hat{W} = \begin{pmatrix} W \\ \zeta \end{pmatrix} = \begin{pmatrix} \beta \\ w \\ \zeta \end{pmatrix},$$

where W is **Planck's temperature vector**
and ζ is **Planck's potential**

- The transformation law $\hat{W}' = \hat{P}^{-1} \hat{W}$ leads to the **normal form**

$$\hat{W} = \begin{pmatrix} \beta \\ \beta v \\ \zeta_{int} + \frac{\beta}{2} \| v \|^2 \end{pmatrix}.$$

Friction tensor

Friction tensor

The **friction tensor** is a linear map from $T_x\mathcal{M}$ into itself, hence a **mixed tensor f** of rank 2

$$f = \nabla \vec{W}$$

represented by the 4×4 matrix $f = \nabla W$

$$= \begin{pmatrix} \frac{\partial \beta}{\partial t} & \frac{\partial \beta}{\partial x} \\ \frac{\partial}{\partial t}(\beta v) - \beta g + \Omega \times \beta v & \frac{\partial}{\partial x}(\beta v) + \beta j(\Omega) \end{pmatrix}$$

- This object introduced by Souriau merges the **temperature gradient** and the **strain velocity**
- In dimension 5, we can also introduce $\hat{f} = \nabla \hat{W}$
represented by a 5×4 matrix $\hat{f} = \nabla \hat{W} = \begin{pmatrix} f \\ \nabla \zeta \end{pmatrix}$

Dual of the friction : momentum tensor

Momentum tensor

Linear map from $T_{\hat{f}(x)}\hat{\mathcal{M}}$ into $T_x\mathcal{M}$, hence a **mixed tensor** \hat{T} of rank 2

it is represented by a 4×5 matrix of the form $\hat{T} = \begin{pmatrix} \mathcal{H} & -p^T & \rho \\ k & \sigma_* & p \end{pmatrix}$

- ρ as the **density**
- $p = \rho v$ as the **linear momentum**
- $\sigma_* = \sigma - \rho v v^T$ as the **dynamical stresses**
- $\mathcal{H} = \rho (e_{int} + \frac{1}{2} \|v\|^2)$ as the **total energy**
- $k = h + \mathcal{H}v - \sigma v$ as the **energy flux** by
conduction convection stress

with :

- the **heat flux** $h = R h'$
- the **statical stresses** $\sigma = R \sigma' R^T$

Friction and momentum tensor

Theorem : reversible medium

if ζ is a function of

- the right Cauchy strain $C = F^T F$
- the temperature vector W
- and the Lagrangean coordinates x_0

then the 4×5 matrix $\hat{T}_R = (T_R, \rho U)$ with $T_R = U \otimes \Pi_R + \begin{pmatrix} 0 & 0 \\ -\sigma_R v & \sigma_R \end{pmatrix}$

$$\text{where } \Pi_R = -\rho \frac{\partial \zeta}{\partial W} \quad \sigma_R = -\frac{2\rho}{\beta} F \frac{\partial \zeta}{\partial C} F^T$$

is a momentum tensor such that :

◊ $Tr(\hat{T}_R \nabla \hat{W}) = 0$ (justifies the terminology of friction tensor)

Second principle

- **Additive decomposition of the momentum tensor** $\hat{T} = \hat{T}_R + \hat{T}_I$
- **Thermodynamics of irreversible processes (TIP)** : Planck's potential depends on extra internal variables (for instance the plastic strain F^P)

Covariant form of the second principle

The **local production of entropy** of a medium characterized by a temperature vector \hat{W} and a momentum tensor \hat{T} is non negative

$$\Phi = \text{Div} \left(\hat{T} \hat{W} \right) - \left(\tau(f(\vec{U})) \right) \left(\tau(T_I(\vec{U})) \right) \geq 0$$

and vanishes if and only if the process is reversible

[de Saxcé & Vallée IJES 2012]

- In the classical form, we recover **Clausius-Duhem inequality**

$$\Phi = \rho \frac{ds}{dt} - \frac{\rho}{\theta} \frac{dq_I}{dt} + \text{div} \left(\frac{h}{\theta} \right) \geq 0$$

Dissipative constitutive laws

Theorem

If $\text{Div } \hat{\mathbf{T}} = \mathbf{0}$ (First principle), the local production of entropy reads

$$\Phi = h \cdot \text{grad } \beta + \beta \text{ Tr}(\sigma_I D)$$

that reveals the duality between

thermodynamic forces (or affinities) $a = \text{grad } \beta$, $A = \beta \text{ grad}_s v = \beta D$
and corresponding **thermodynamic fluxes** h, σ_I

Assuming a linear isotropic law $(h, \sigma_I) = F(a, A)$,
we recover the heat conduction and Navier-Stokes equations

Epistemological reversal : coming back to the relativistic model

Poincaré general relativity

Relativistic form of the 2nd principle

The **local production of entropy** of a medium characterized by a temperature vector \vec{W} and a momentum tensor T is non negative :

$$\Phi = \mathbf{Div} (\mathbf{T} \vec{W} + \zeta \vec{N}) - \frac{1}{c^2} \left(\mathbf{U}^*(\mathbf{f}(\vec{U})) \right) \frac{1}{c^2} \left(\mathbf{U}^*(\mathbf{T}_I(\vec{U})) \right) \geq 0 ,$$

and vanishes if and only if the process is reversible

- **Dissipative constitutive laws** : see Claude Vallée, IJES (1981)

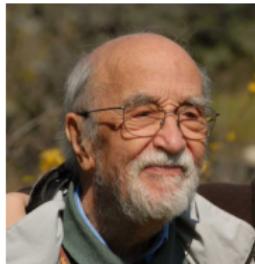
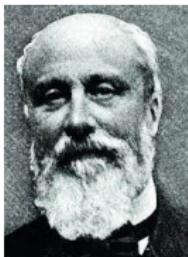
In contrast to the Galilean case, there is a metric

The momentum tensor T is self-adjoint and the friction too :

$$\mathbf{f} = \frac{1}{2} \left[\nabla \vec{W} + (\nabla \vec{W})^* \right]$$

Thank you !

This work is dedicated to Pierre Duhem



**66th SOURIAU
COLLOQUIUM
(CITV)
Bastia (Corse)
28 avril au 3 mai
2024**

