

# Développement asymptotique de la relativité générale avec covariance galiléenne

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# Getting started

## As model, we take the Relativity

- The space-time  $\mathcal{M}$  is **Riemannian** with **10** potentials  $G_{\alpha\beta}$ , solutions of the **10 Einstein field equations** :

$$R'_{\alpha\beta} - \frac{1}{2} R G_{\alpha\beta} + \Lambda G_{\alpha\beta} = \frac{8\pi k_g}{c^4} T_{\alpha\beta}$$

## Is this scheme transposable to classical mechanics ?

- We consider **Galileo** symmetry group :  
 $\mathcal{M}$  is not **Riemannian** !  
 but **there is an integrable G-structure**
- The associated connection (Galilean gravitation) has 2 components  $g, \Omega$   
 deriving from **4** potentials  $\phi, A^i$
- Which are the **4** corresponding **field equations** ?

# State of the Art

- **Post-Newtonian** theory [Eddington 1922, Will 1971, Ni 1972, Weinberg 1972, Misner 1973, ...]
- **Newton-Cartan** theory [Cartan 1923, Trautman 1963, Souriau 1970, Küntzle 1972, Duval 1985, Horváthy 1991, Bergshoeff 2011, ...]

Trautmann : Ricci tensor is :  $R' = -\text{div } g \tau \otimes \tau$

**flaw** : as Poisson's equation  $-\text{div } g = \Delta \phi = 4\pi k_g \rho$ ,  
it has not the expected Galilean covariance

- **Merging both approaches** [Dautcourt 1997, Tichy & Flanagan 2011, Van den Bleeken 2017]

**Present approach** : based on a variational version  
of General Relativity and the use of the  $G$ -structure

[GERG de Saxcé 2020 Asymptotic expansion of GR  
with Galilean covariance ]

## Galileo's group and geometry

- **Chart** :  $\mathbb{R}^4 \ni X = \begin{pmatrix} t \\ x \end{pmatrix} \mapsto \mathbf{X} \in \mathcal{M}$  space-time
- **Galileo's group**  $\text{GAL}$  is the one of the **Galilean transformations** :

$$X = P X' + C \quad \text{with} \quad P = \begin{pmatrix} 1 & 0 \\ u & R \end{pmatrix}, \quad C = \begin{pmatrix} \tau_0 \\ k \end{pmatrix}$$

where  $u \in \mathbb{R}^3$  is the **Galilean boost** or **velocity of transport** and  $R$  is a rotation

- $\text{GAL}_0$  is the subgroup of *linear* Galilean transformation  $P$

## Galileo's group and geometry

## Theorem

$X \mapsto X'$  is such that  $\frac{\partial X'}{\partial X} = P^{-1}$  is Galilean  
 iff  $x' = (R(t))^T (x - x_0(t)), \quad t' = t + \tau_0$

- **Proof** : find the compatibility conditions of  $\frac{\partial X'}{\partial X} = P^{-1}$  (Frobenius)
- This defines an **equivalence class** of coordinate charts  
 $X \mapsto X'$  is called a **galileomorphism**,  
 $X$  and  $X'$  are called **Galilean coordinate systems** or **Galilean charts**  
 In a physical point of view, they are **the frames in which the durations and distances are measured**
- Example of **integrable  $G$ -structure**  $L_G$  [Kobayashi 1963]

## Galileo's group and geometry

## Theorem

A **Galilean gravitation** is a symmetric connection  $\Gamma$  which, in a Galilean chart, takes its values in the Lie algebra  $\mathfrak{gal}_0$  of the group of linear Galilean transformations :

$$\Gamma(dX) = \begin{pmatrix} 0 & 0 \\ \Omega \times dx - g dt & j(\Omega) dt \end{pmatrix},$$

where  $j(\Omega)$  is the unique skew-symmetric matrix such that  $j(\Omega)v = \Omega \times v$

- $g$  is the classical **gravity**
- $\Omega$  is a new object called **spinning**

## Galileo's group and geometry

4-velocity  $U = \dot{X} = \begin{pmatrix} 1 \\ v \end{pmatrix}$  and linear 4-momentum  $T = m U$

## Covariant law of the motion

For particles in the gravitation field

$$\nabla_U T = \dot{T} + \Gamma(U) T = 0$$

in a Galilean chart :  $\dot{m} = 0, \quad \dot{p} = m(g - 2\Omega \times v)$

[Souriau 1970 Structure des systèmes dynamiques]

The **inertial charts** are the ones in which the spinning  $\Omega$  vanishes

## Galileo's group and geometry

- The closure condition  $d\omega = 0$  of the presymplectic form gives

$$\text{curl } g + 2 \frac{\partial \Omega}{\partial t} = 0, \quad \text{div } \Omega = 0$$

then there exist **potentials**  $\phi, A$  such that

$$g = -\text{grad } \phi - \frac{\partial A}{\partial t}, \quad \Omega = \frac{1}{2} \text{curl } A$$

- They are defined modulo a **gauge** :  $\phi^* = \phi - \frac{\partial f}{\partial t}$ ,  $A^* = A + \text{grad } f$
- The corresponding Lagrangian  $\mathcal{L} = \frac{1}{2} m \|v\|^2 + m A \cdot v - m \phi$  is covariant provided

$$\phi' = \phi - A \cdot u - \frac{1}{2} \|u\|^2, \quad A' = R^T(A + u)$$



# Galilean tensors

- By restriction of the transformation law of tensors to a subgroup  $G \subset \text{GL}(n)$ , we obtain the  **$G$ -tensors**
  - Euclidean tensors are  $\text{SO}(3)$ -tensors
  - Galilean tensors are  $\text{GAL}_0$ -tensors
- $G$ -tensors may be seen as **orbits** for the action of  $G$  onto the tensor components
- Examples of Galilean tensors :
  - **Vectors** :  $V^{\alpha'} = (P^{-1})^{\alpha'}_{\beta} V^{\beta}$  or  $V' = P^{-1}V$  with  $P \in \text{GAL}_0$   
 applied to  $U = \begin{pmatrix} 1 \\ v \end{pmatrix}$  gives  $v' = R^T(v - u)$
  - **Covectors** :  $\Phi_{\alpha'} = \Phi^{\beta} P^{\beta}_{\alpha'}$ , or  $\Phi' = \Phi P$  applied to  $\tau = (1, 0^T)$   
 invariant representing the clock form  $\tau = dt$  (**absolute time**)

# Galilean tensors

Examples of Galilean tensors :

- **Vectors**
- **Covectors**
- **2-covariant tensors** :  $T_{\alpha'\beta'} = P_{\alpha'}^{\mu} P_{\beta'}^{\nu} T_{\mu\nu}$  or  $T' = P^T T P$

applied to  $T = \begin{pmatrix} a & w^T \\ w & M \end{pmatrix}$  symmetric gives

$$a' = a + 2w \cdot u + u(Mu), \quad w' = R^T(w + Mu), \quad M' = R^T M R$$

For instance  $T = \overset{(0)}{G} = \begin{pmatrix} 2\phi & -A^T \\ -A & -1_{\mathbb{R}^3} \end{pmatrix}$  is a Galilean 2-covariant tensor  
(but is not a metrics!)

# Spacetime metrics

For weak gravitational fields

- $G = \begin{pmatrix} c^2 + 2\phi & -A^T \\ -A & -1_{\mathbb{R}^3} \end{pmatrix} = \epsilon^{-1} \overset{(-1)}{G} + \overset{(0)}{G}$  (exact, not a, expansion !)

where  $\epsilon = c^{-2}$ ,  $\overset{(-1)}{G} = \tau \otimes \tau$  and  $\overset{(0)}{G}$  are Galilean tensors

- Truncated expansion of

$$G^{-1} \cong \begin{pmatrix} 0 & 0 \\ 0 & -1_{\mathbb{R}^3} \end{pmatrix} + \epsilon \begin{pmatrix} 1 & -A^T \\ -A & AA^T \end{pmatrix}$$

Each term is a Galilean 2-contravariant tensor

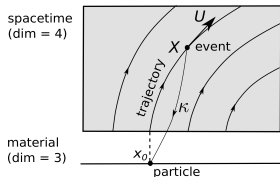
- $\sqrt{-\det G} \cong c \left( 1 + \epsilon \left( \phi + \frac{\|A\|^2}{2} \right) \right) = c (1 + \epsilon l_0)$

Galilean invariant

Hilbert-Einstein functional  $\mathcal{A} = \int (p_M + p_G) \sqrt{-\det G} d^4 X$

## Functional due to the matter

Its motion is described  
by a line bundle  
of projection

$$x_0 = \kappa(X) = \kappa(t, x)$$


- **Conformation tensor** [Souriau 1964 Géométrie & Relativité]

$$\mathcal{D} = -\frac{\partial x_0}{\partial X} \left( \frac{\partial x_0}{\partial X} \right)^* = -\frac{\partial x_0}{\partial X} G^{-1} \left( \frac{\partial x_0}{\partial X} \right)^T$$

where  $\frac{\partial x_0}{\partial X} = F^{-1}(-v, 1_{\mathbb{R}^3})$  with  $F = \frac{\partial x}{\partial x_0}$

- mass density  $\rho = \rho_0(x_0) / \det F$  which is a Galilean invariant

$$p_M \sqrt{-\det G} = c^2 \rho_0(x_0) \sqrt{\det \mathcal{D} (-\det G)} \cong c \left[ \rho c^2 - \rho \left( \frac{\|v\|^2}{2} + A \cdot v - \phi \right) \right]$$

energy

Lagrangian

at rest

density

# Curvature tensor

We are working now with two structures

- **The Galilean structure** : we work in the charts of the  $G$ -structure
- **The Riemannian structure** : we equip the manifold with a metrics

**Both structures are compatible because we use an asymptotic expansion of the metrics of which the terms are Galilean tensors**

## Curvature tensor

Levi-Civita connection  $\Gamma(dX) = (\Gamma_{\alpha\beta}^{\mu} dX^{\alpha}) \cong \overset{(0)}{\Gamma} + \epsilon \overset{(1)}{\Gamma}$

- **Order 0** (Galilean gravitation)

$$\overset{(0)}{\Gamma}_{00}^i = -g^i \quad \text{gravity}$$

$$\overset{(0)}{\Gamma}_{i0}^j = \Omega_i^j \quad \text{spinning}$$

- **Order 1**

$$a = \overset{(1)}{\Gamma}_{00}^0 = \frac{\partial \phi}{\partial t} - A \cdot g$$

$$B_i = \overset{(1)}{\Gamma}_{0i}^0 = \delta_{ik} (grad \phi - \Omega \times A)^k$$

$$D_{ij} = \overset{(1)}{\Gamma}_{ij}^0 = -\delta_{ik} (grad_s A)_j^k$$

May be also obtained from the Bargmannian connection :  $\overset{(1)}{\Gamma}_{\alpha\beta}^0 = \overset{(1)}{\Gamma}_{\alpha\beta}^4$

## Curvature tensor

Ricci tensor  $R'_{\beta\gamma} = R^{\alpha}_{\alpha\beta\gamma}$   
 expanded as  $R' \cong R^{(0)'} + \epsilon R^{(1)'}$

- Order 0

$$R^{(0)'} = \begin{pmatrix} 2 \|\Omega\|^2 - \text{div } g & (\text{curl } \Omega)^T \\ \text{curl } \Omega & 0 \end{pmatrix}$$

- Order 1 for instance :

$$R^{(1)'}_{lm} = D_{lj}\Omega_m^j + D_{jm}\Omega_l^j + \frac{\partial}{\partial x^j}(A^j D_{lm}) - \frac{\partial D_{lm}}{\partial t} - \frac{\partial}{\partial x^l}(B_m + D_{jm}A^j)$$

...

As expected, symmetric because  $\frac{\partial}{\partial x^j}(B + D A)$  is the Hessian of  $I_0$

# Functional due to the geometry

$$p_G \sqrt{-\det G} = -\frac{c^4}{8\pi k_g} (2\Lambda\epsilon + R) \sqrt{-\det G},$$

where the **scalar curvature** is  $R = G^{\alpha\beta} R'_{\alpha\beta} = \text{Tr}(G^{-1}R')$

that leads to  $R \cong -2\epsilon(I + I_1)$  with the **Galilean invariants** of

- the metrics and the curvature :  $I = \text{div } g - 2 \|\Omega\|^2 + 2A \cdot \text{curl } \Omega$
- only the curvature :  $I_1 = -\frac{1}{2} \text{Tr} [(D - (\text{Tr } D)1_{\mathbb{R}^3})^2]$



# Functional due to the geometry

## Complete functional

$$c \int \left[ \frac{\epsilon^{-1}}{4\pi k_g} (I + I_1 - \Lambda) (1 + \epsilon I_0) + \rho \epsilon^{-1} - \rho \left( \frac{\|v\|^2}{2} + A \cdot v - \phi \right) \right] d^4 X$$

where  $\overbrace{\text{div } g - 2 \|\Omega\|^2 + 2 A \cdot \text{curl } \Omega}$   $\overbrace{\left( \phi + \frac{\|A\|^2}{2} \right)}$

## 4 field equations (not 10!)

- Variation wrt  $\phi$  :  $I + I_1 - \Lambda = -4\pi k_g \rho$
- Variation wrt  $A$  :  $-2(\epsilon^{-1} + I_0) \text{curl } \Omega - (I + I_1 - \Lambda) A = -4\pi k_g \rho v$

## Consistency of the equations :

- The rhs  $-4\pi k_g \rho U$  is a Galilean vector (because  $\rho$  is invariant)
- The lhs is also a Galilean vector

## Asymptotic expansion of the solution

Reminder :

$$\text{Variation wrt } \phi : I + I_1 - \Lambda = -4\pi k_g \rho \quad (t)$$

$$\text{Variation wrt } A : -2(\epsilon^{-1} + I_0) \text{curl } \Omega - (I + I_1 - \Lambda) A = -4\pi k_g \rho v \quad (s)$$

**Solution in 3 steps :**

① (s) at order  $-1$  :  $\text{curl } \overset{(0)}{\Omega} = 0$

remember :  $\text{div } \overset{(0)}{\Omega} = 0$ , then  $\overset{(0)}{\Omega}$  is harmonic

**hypothesis** :  $\overset{(0)}{\Omega}$  is bounded, then  $\overset{(0)}{\Omega} = \overset{(0)}{\Omega}(t)$

**in agreement with observations but justified by GR**

Next  $\overset{(0)}{A} = -u$  then  $\overset{(0)}{D} = 0$  and  $\overset{(0)}{g} = -\text{grad } \overset{(0)}{\phi}$

**where occurs the velocity of transport**

## Asymptotic expansion of the solution

$$\text{Variation wrt } \phi : I + I_1 - \Lambda = -4\pi k_g \rho \quad (t)$$

$$\text{Variation wrt } A : -2(\epsilon^{-1} + I_0) \text{curl } \Omega - (I + I_1 - \Lambda) A = -4\pi k_g \rho v \quad (s)$$

$$\textcircled{1} \text{ (s) at order } -1 : \text{curl } \overset{(0)}{\Omega} = 0 \text{ then } \overset{(0)}{\Omega} = \overset{(0)}{\Omega}(t)$$

$$\text{Next } \overset{(0)}{A} = -u \text{ then } \overset{(0)}{D} = 0 \text{ and } \overset{(0)}{g} = -\text{grad } \overset{(0)}{\phi}$$

## Asymptotic expansion of the solution

$$\text{Variation wrt } \phi : I + I_1 - \Lambda = -4\pi k_g \rho \quad (t)$$

$$\text{Variation wrt } A : -2(\epsilon^{-1} + I_0) \text{curl } \Omega - (I + I_1 - \Lambda) A = -4\pi k_g \rho v \quad (s)$$

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$$\text{Next } \overset{(0)}{A} = -u \text{ then } \overset{(0)}{D} = 0 \text{ and } \overset{(0)}{g} = -\text{grad } \overset{(0)}{\phi}$$

$$\textcircled{2} \text{ (t) at order } 0 : I + I_1 = -\Delta \overset{(0)}{\phi} - 2 \|\overset{(0)}{\Omega}\|^2 = -4\pi k_N \rho + \Lambda$$

$$\overset{(0)}{\phi} = -\int \frac{k_N \rho(x', t)}{\|x - x'\|} dV(x') - \frac{1}{2} \|u\|^2 - \frac{\Lambda}{6} \|x\|^2$$

## Asymptotic expansion of the solution

$$\text{Variation wrt } \phi : I + I_1 - \Lambda = -4\pi k_g \rho \quad (t)$$

$$\text{Variation wrt } A : -2(\epsilon^{-1} + I_0) \text{curl } \Omega - (I + I_1 - \Lambda) A = -4\pi k_g \rho v \quad (s)$$

$$\textcircled{1} \text{ (s) at order } -1 : \text{curl } \overset{(0)}{\Omega} = 0 \text{ then } \overset{(0)}{\Omega} = \overset{(0)}{\Omega}(t)$$

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$$\textcircled{3} \text{ (s) at order } 0 : -2 \text{curl } \overset{(1)}{\Omega} = -\text{grad}(\text{div } \overset{(1)}{A}) + \Delta \overset{(1)}{A} = -4\pi k_N \rho v$$

choice of a gauge st  $= 0$

$$\overset{(1)}{A} = \int \frac{k_N \rho(x', t) v(x', t)}{\|x - x'\|} dV(x')$$

relativistic effect

# Conclusions and perspectives

## Conclusions

- The scheme of the General Relativity is transposable to the classical mechanics. Instead of expressing the identity between two tensors of rank 2, the field equations identify two 4-vectors. **They allow to determine not only the gravity but also the spinning**
- Inspired from both Newton-Cartan and Post-Newtonian theories, the present approach leads to covariant field equations given in the charts of the  $G$ -structure, that is **the frames in which the durations and distances are measured**

# Conclusions and perspectives

## Perspectives

- **Closed analytical and numerical solutions**

of the (non linear) field equations

$$g = -grad \phi - \frac{\partial A}{\partial t}, \quad \Omega = \frac{1}{2} curl A, \quad I_1 = -\frac{1}{2} Tr [(D - (Tr D) 1_{\mathbb{R}^3})^2]$$

$$div g - 2 \|\Omega\|^2 + 2 A \cdot curl \Omega + I_1 + \Lambda = -4 \pi k_N \rho$$

$$-2 \left( c^2 + \phi + \frac{\|A\|^2}{2} \right) curl \Omega$$

$$-(div g - 2 \|\Omega\|^2 + 2 A \cdot curl \Omega + I_1 + \Lambda) A = -4 \pi k_N \rho v$$

- **Extension to order 2**
- **Towards the General Relativity**

with 5 potentials  $G = \begin{pmatrix} c^2 + 2\phi & -A^T \\ -A & -\psi 1_{\mathbb{R}^3} \end{pmatrix}$

with 10 potentials  $G = \begin{pmatrix} c^2 + 2\phi & -A^T \\ -A & -M \end{pmatrix}$

# Thank you for your attention !

Any questions ?