Principle of General Covariance : Applications to Inelastic Continuum and Relative Gravitation

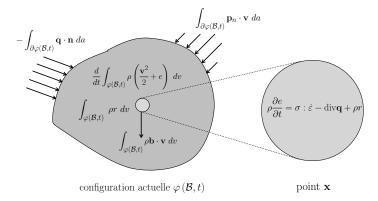
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- Classical Conservations Laws
- **2** Relative Gravitation and Principle of General Covariance
- **③** Description of some Inelastic Deformation
- **9** Principle of General Covariance and Applications
- Oncluding remarks

I. Conservation laws : Classical Thermomechanics



Conservation of energy in global formulation and localization.

Conservation laws : unified framework

B-derivative (particular Lie derivative) (R 2003, 2022)

$$\left(\frac{d^{\mathscr{B}}}{dt}\mathcal{T}\right)\left(\mathbf{u}_{1},...\mathbf{u}_{p},\omega^{1},...\omega^{q}\right):=\frac{d}{dt}\left[\mathcal{T}\left(\mathbf{u}_{1},...\mathbf{u}_{p},\omega^{1},...\omega^{q}\right)\right]$$

2 Conservation laws (Poincaré's integral invariance) not covariant

$$\frac{d^{\mathscr{B}}}{dt}(\rho\omega_{n}) = 0$$

$$\frac{d^{\mathscr{B}}}{dt}(\rho\mathbf{v}^{i}\omega_{n}) = \rho b^{i}\omega_{n} + \operatorname{div}\mathbf{p}_{n}^{i}\omega_{n}, \quad i = 1, 2, 3$$

$$\frac{d^{\mathscr{B}}}{dt}\left[\rho\left(u + \frac{\mathbf{v}^{2}}{2}\right)\omega_{n}\right] = \rho \mathbf{b} \cdot \mathbf{v} \,\omega_{n} + \operatorname{div}\left(\sigma^{\mathrm{T}}(\mathbf{v})\right)\omega_{n} + \rho r \,\omega_{n} - \operatorname{div}\mathbf{J}_{H} \,\omega_{n}$$

Intropy inequality

$$\frac{d^{\mathscr{B}}}{dt}\left(\rho s\omega_{n}\right) \geq \rho \frac{r}{\theta} \, \omega_{n} - \operatorname{div}\left(\frac{\mathbf{J}_{H}}{\theta}\right)\omega_{n}$$

Conservation laws : embedded base $(\mathbf{u}_{\alpha}, \alpha = 1, 3)$

Mass

$$\rho_0 = \rho \ J, \quad \text{with} \quad J := \frac{\omega_n \left(\mathbf{u}_1, \mathbf{u}_2, \mathbf{u}_3 \right)}{\omega_n \left(\mathbf{u}_{10}, \mathbf{u}_{20}, \mathbf{u}_{30} \right)}$$

2 Linear momentum:

$$\rho_0 \frac{d}{dt} \mathbf{v} = \rho_0 \mathbf{b} + \text{Div} \mathbf{P}$$

S Energy
$$\rho_0 \frac{d}{dt} u = \mathbf{P} : \nabla \mathbf{v} + \rho_0 r - \text{Div} \mathbf{J}_{H0}$$
S Entropy
$$\rho_0 \frac{d}{dt} \mathbf{s} \ge \rho_0 \frac{\mathbf{r}}{\theta} - \text{Div} \left(\frac{\mathbf{J}_{H0}}{\theta} \right)$$

Conservation laws : "spatial" base $(\mathbf{e}_i, i = 1, 3)$

Mass

$$\frac{\partial \rho}{\partial t} + \operatorname{div}\left(\rho \mathbf{v}\right) = \mathbf{0}$$

2 Linear momentum:

$$\frac{\partial}{\partial t} \left(\rho \mathbf{v} \right) + \operatorname{div} \left(\rho \mathbf{v} \otimes \mathbf{v} \right) = \rho \, \mathbf{b} + \operatorname{div} \sigma$$

$$\rho \frac{d}{dt} u = \sigma : \mathbf{D} + \rho r - \operatorname{div} \mathbf{J}_H$$

Intropy

$$\rho \frac{d}{dt} s \ge \rho \frac{r}{\theta} - \operatorname{div}\left(\frac{\mathbf{J}_H}{\theta}\right)$$

Remarque

Base (\mathbf{e}_i , i = 1, 3) is temporarily embedded base within the continuum at t (Goldstein "Lecture Notes on Fluid Mechanics").

Symmetries and conservation laws

- Despite their unique framework for expressing conservation laws, Newton's laws of motion are not covariant (Galilean transformations) conversely to Lagrange's equations.
- Einstein's great advance was to put spacetime symmetry first : SR makes particle mechanics and electromagnetism compatible (Lorentz transformations)
- For RG, Einstein introduced the principle of equivalence (local symmetry) invariance of law expressions under local changes of the spacetime coordinates. Symmetry Invariance for RG requires additional precisions: it was a source of disputes (Norton, 2005).
- Indeed, Noether's theorems relate symmetry with conservation laws: continuous symmetries generate conservation laws.
- Constructing conservation laws and conserved quantities in an physics theory, including Einstein-Cartan Relative Gravitation and also other field theory, is still a main problem.

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II. Relative Gravitation and Principle of General Covariance



- Spacetime geometry
- Relative Gravitation as Gauge Theory

Tool 1. : Connection, Torsion and Curvature

Given an affine **connection** ∇ on a manifold \mathscr{B} :

Constants of structure of Cartan (non coordinate base)¹

$$[\mathbf{e}_{\alpha},\mathbf{e}_{\beta}] := \aleph_{\mathbf{0}\alpha\beta}^{\gamma} \ \mathbf{e}_{\gamma}$$

② Torsion tensor

$$\begin{cases} \aleph(\mathbf{f}^{\gamma}, \mathbf{e}_{\alpha}, \mathbf{e}_{\beta}) = \mathbf{f}^{\gamma} \left(\nabla_{\mathbf{e}_{\alpha}} \mathbf{e}_{\beta} - \nabla_{\mathbf{e}_{\beta}} \mathbf{e}_{\alpha} - [\mathbf{e}_{\alpha}, \mathbf{e}_{\beta}] \right) \\ \aleph_{\alpha\beta}^{\gamma} = \Gamma_{\alpha\beta}^{\gamma} - \Gamma_{\beta\alpha}^{\gamma} - \aleph_{0\alpha\beta}^{\gamma} \end{cases}$$

Ourvature tensor

$$\begin{cases} \Re(\mathbf{f}^{\gamma}, \mathbf{e}_{\alpha}, \mathbf{e}_{\beta}, \mathbf{e}_{\lambda}) = \mathbf{f}^{\gamma}(\nabla_{\mathbf{e}_{\alpha}} \nabla_{\mathbf{e}_{\beta}} \mathbf{e}_{\lambda} - \nabla_{\mathbf{e}_{\beta}} \nabla_{\mathbf{e}_{\alpha}} \mathbf{e}_{\lambda} - \nabla_{[\mathbf{e}_{\alpha}, \mathbf{e}_{\beta}]} \mathbf{e}_{\lambda}) \\ \Re^{\gamma}_{\alpha\beta\lambda} = (\partial_{\alpha} \Gamma^{\gamma}_{\beta\lambda} + \Gamma^{\mu}_{\beta I} \Gamma^{\gamma}_{\alpha\mu}) - (\partial_{\beta} \Gamma^{\gamma}_{\alpha\lambda} + \Gamma^{\mu}_{\alpha\lambda} \Gamma^{\gamma}_{\beta\mu}) - \aleph^{\mu}_{0\alpha\beta} \Gamma^{\gamma}_{\mu\lambda} \end{cases}$$

¹Orthonormal cylindrical : $(\mathbf{e}_r, \mathbf{e}_\theta, \mathbf{e}_z)$ and spherical $(\mathbf{e}_r, \mathbf{e}_\theta, \mathbf{e}_\varphi)$ are not coordinate bases.

Tool 2. : Covariant differential operators for \mathscr{B}

Two covariant differential operators necessary for any continuum theory. Let **T** be a tensor type (p, q):

Connection and covariant derivative ON B

$$\overline{\nabla_{\gamma} T^{\alpha_{1} \cdots \alpha_{p}}_{\beta_{1} \cdots \beta_{q}}} = \partial_{\gamma} T^{\alpha_{1} \cdots \alpha_{p}}_{\beta_{1} \cdots \beta_{q}} + \sum_{s=1}^{s=p} \Gamma^{\alpha_{s}}_{\gamma \mu} T^{\alpha_{1} \cdots \alpha_{s-1} \ \mu \ \alpha_{s+1} \cdots \alpha_{p}}_{\beta_{1} \cdots \beta_{q}} - \sum_{s=1}^{s=q} \Gamma^{\mu}_{\gamma \beta_{s}} T^{\alpha_{1} \cdots \alpha_{p}}_{\beta_{1} \cdots \beta_{s-1} \ \mu \ \beta_{s+1} \cdots \beta_{q}}$$

2 Lie derivative generates an active diffeomorphism OF *B*

$$\begin{array}{lll} \underbrace{\mathcal{L}_{\xi} T^{\alpha_{1} \cdots \alpha_{p}}_{\beta_{1} \cdots \beta_{q}}}_{+ & T^{\alpha_{1} \cdots \alpha_{p}}_{\beta_{1} \cdots \beta_{q}} \end{array} & := & \xi^{\gamma} \partial_{\gamma} T^{\alpha_{1} \cdots \alpha_{p}}_{\beta_{1} \cdots \beta_{q}} \\ & - & T^{\gamma \alpha_{2} \cdots \alpha_{p}}_{\beta_{1} \cdots \beta_{q}} & \partial_{\gamma} \xi^{\alpha_{1}} - \cdots - T^{\alpha_{1} \cdots \alpha_{p-1} \gamma}_{\beta_{1} \cdots \beta_{q}} & \partial_{\gamma} \xi^{\alpha_{p}} \\ & + & T^{\alpha_{1} \cdots \alpha_{p}}_{\gamma \beta_{2} \cdots \beta_{q}} & \partial_{\beta_{1}} \xi^{\gamma} + \cdots + T^{\alpha_{1} \cdots \alpha_{p}}_{\beta_{1} \cdots \beta_{q-1} \gamma} & \partial_{\beta_{q}} \xi^{\gamma} \end{array}$$

Tool 3. : Derivative of an integral on \mathscr{B}

1 Definition of a **volume-form** (measure) and action on \mathscr{B}

$$\omega_n, \qquad \mathscr{S} := \int_{\mathscr{B}} \mathscr{L} \, \omega_n$$

where \mathscr{L} is a Lagrangian density of a action.

Poincaré: Variation of integral under active diffeomorphism. Let ω be a p-form (p ≤ n) on ℬ. Then under the flow (with tangent ξ):

$$\int_{\mathscr{B}} \omega \quad \text{integral invariant} \quad \Longleftrightarrow \quad \boxed{\mathcal{L}_{\xi} \ \omega = \mathbf{0}}$$

 Application on an action (remark presence of the second term ! e.g. Obukhov & Putzfield, 2014):

$$\mathscr{S} = \int_{\mathscr{B}} \mathscr{L}\omega_n \quad \Longrightarrow \quad \delta\mathscr{S} = \int_{\mathscr{B}} \mathcal{L}_{\xi}\left(\mathscr{L}\right)\omega_n + \left[\int_{\mathscr{B}} \mathscr{L} \mathcal{L}_{\xi}\omega_n\right]$$

Continuum 1. : Riemann–Cartan Continuum

Continuum is a compact, connected, oriented manifold :

- a metric tensor **g** with components $g_{lphaeta}(\mathbf{x})$
- an affine connection ∇ with coefficients $\Gamma^{\gamma}_{\alpha\beta}(\mathbf{x})$
- a volume-form ω_n with components $\omega_{n \ 1 \cdots n}(\mathbf{x})$

2 ∇ and **g** are **compatible** if

$$abla_{\gamma} g_{lpha eta} \equiv 0$$

• The second state of the

$$\mathcal{L}_{\xi}\omega_{n}=\left(\nabla_{\alpha}\xi^{\alpha}\right)\omega_{n}$$

 ξ tangent to flow (active diffeomorphism). ω_n exists !

Continuum 2. : Riemann Continuum

Continuum is a compact, connected, oriented manifold :

- a metric tensor **g** with components $g_{lphaeta}\left(\mathbf{x}
 ight)$
- a Levi-Civita connection $\overline{\nabla}$ with coefficients $\overline{\Gamma}_{\alpha\beta}^{\gamma}(\mathbf{x})$
- a volume-form $\overline{\omega}_n := \sqrt{\operatorname{Det} \mathbf{g}} \ dx^1 \wedge \cdots \wedge dx^n$

2 Symmetric connection $\overline{\nabla}$ and metric g are compatible

$$\overline{
abla}_{\gamma} g_{lphaeta} \equiv \mathsf{0}, \qquad \overline{\mathsf{\Gamma}}_{lphaeta}^{\gamma} = (1/2) \; g^{\gamma\lambda} \left(\partial_{eta} g_{lpha\lambda} + \partial_{lpha} g_{\lambdaeta} - \partial_{\lambda} g_{lphaeta}
ight)$$

3 Symmetric connection $\overline{\nabla}$ and $\overline{\omega}_n$ are **compatible**

$$\mathcal{L}_{\xi}\overline{\omega}_{n} = \left(\overline{\nabla}_{\alpha}\xi^{\alpha}\right)\overline{\omega}_{n}$$

 ξ tangent to flow (active diffeomorphism). (**Divergence theorem**)

Relative Gravitation : Einstein spacetime (Riemann continuum)



• The **Hilbert-Einstein** action governs the (simplest) theory for **vacuum spacetime** in Relative Gravitation :

$$\mathscr{S} := \frac{1}{2\chi} \int_{\mathscr{M}} \overline{\mathfrak{R}} \, \overline{\omega}_n, \quad \overline{\mathfrak{R}} := g_{\alpha\beta} \, \overline{\mathfrak{R}}^{\alpha\beta}, \quad \overline{\mathfrak{R}}_{\alpha\beta} := \mathfrak{R}^{\gamma}_{\gamma\alpha\beta}$$

1. Gravitation : Two-steps approach e.g. Ali et al. 2009

1 The **Euler variation** $\delta g_{\alpha\beta}$ induces the Einstein field equation:

$$\Delta \mathscr{S} = (1/2\chi) \int_{\mathscr{M}} \left(\overline{\mathfrak{R}}^{\alpha\beta} - (\overline{\mathfrak{R}}/2) g^{\alpha\beta} \right) \delta g_{\alpha\beta} \ \overline{\omega}_n + \mathrm{BT} = 0$$

with $\delta\overline{\omega}_n = -(1/2)g^{lphaeta}\,\,\delta g_{lphaeta}\,\,\overline{\omega}_n$

② Einstein field equation (not a conservation laws) holds:

$$\overline{\mathsf{G}}^{\alpha\beta} := \overline{\Re}^{\alpha\beta} - (\overline{\Re}/2)g^{\alpha\beta} = 0$$

Orinciple of General Covariance. Considering active diffeomorphism engendered by Lie derivative along ξ,

$$\delta g_{\alpha\beta} \longrightarrow \mathcal{L}_{\xi} g_{\alpha\beta} = \overline{\nabla}_{\alpha} \xi_{\beta} + \overline{\nabla}_{\beta} \xi_{\alpha}$$

we deduce the vacuum spacetime **conservation laws** :

$$\overline{
abla}_{eta}\overline{G}^{lphaeta}=0$$

2. Gravitation : Palatini approach e.g. Ferraris et al. 1982

U Hilbert-Einstein action (metric, connection and derivatives):

$$\mathscr{S} = (1/2\chi) \int_{\mathscr{B}} \mathscr{L} \left(g_{\alpha\beta}, \overline{\Gamma}^{\gamma}_{\alpha\beta}, \partial_{\lambda} \overline{\Gamma}^{\gamma}_{\alpha\beta} \right) \overline{\omega}_{n}$$

Variations of metric and connection

$$\begin{split} \delta \overline{\Gamma}^{\gamma}_{\alpha\beta} &= (1/2) \ g^{\gamma\lambda} \left(\overline{\nabla}_{\beta} \delta g_{\alpha\lambda} + \overline{\nabla}_{\alpha} \delta g_{\lambda\beta} - \overline{\nabla}_{\lambda} \delta g_{\alpha\beta} \right) \\ \delta \overline{\Re}_{\alpha\beta} &= \overline{\nabla}_{\lambda} (\delta \Gamma^{\lambda}_{\beta\alpha}) - \overline{\nabla}_{\beta} (\delta \Gamma^{\lambda}_{\lambda\alpha}) \end{split}$$

Suler-Lagrange equations as field equations :

$$\begin{cases} G_{\alpha\beta} := \Re_{(\alpha\beta)} \left[\Gamma^{\sigma}_{\mu\nu}, \partial_{\epsilon} \Gamma^{\sigma}_{\mu\nu} \right] - \frac{1}{2} g_{\alpha\beta} \left(g^{\eta\theta} \Re_{(\eta\theta)} \left[\Gamma^{\sigma}_{\mu\nu}, \partial_{\epsilon} \Gamma^{\sigma}_{\mu\nu} \right] \right) = 0 \\ \overline{\nabla}^{\gamma} g_{\alpha\beta} = 0 \end{cases}$$

Remarque

Conservation laws follow from $\delta g_{\alpha\beta} \longrightarrow \mathcal{L}_{\xi} g_{\alpha\beta} = \overline{\nabla}_{\alpha} \xi_{\beta} + \overline{\nabla}_{\beta} \xi_{\alpha}$

3. Gravitation: Integral Invariance & Bianchi's identities

Principle of General Covariance (active diffeomorphism)

$$\delta \mathscr{S} = \int_{\mathscr{B}} \mathcal{L}_{\xi}(\mathscr{L}) \,\overline{\omega}_{n} + \mathscr{L}(\overline{\nabla}_{\alpha}\xi^{\alpha}) \,\overline{\omega}_{n} = \frac{1}{2\chi} \int_{\mathscr{B}} g^{\beta\lambda} \overline{\nabla}_{\nu} \overline{\mathcal{R}}_{\beta\lambda} \,\overline{\omega}_{n} = 0$$

The second term vanishes for **divergence-free gauge** ξ .

Bianchi's first identities

$$\overline{
abla}_{
u}\overline{rak{R}}_{lphaeta\lambda}^{\gamma}+\overline{
abla}_{eta}\overline{rak{R}}_{
ulpha\lambda}^{\gamma}+\overline{
abla}_{lpha}\overline{rak{R}}_{eta
u\lambda}^{\gamma}=0$$

③ Technique : Putting $\alpha = \gamma$ and multiplying by $g^{\beta\lambda}$:

$$\overline{\nabla}_{\nu}\overline{\mathcal{R}} - \overline{\nabla}_{\beta}\overline{\mathcal{R}}_{\nu}^{\beta} - \overline{\nabla}_{\alpha}\overline{\mathcal{R}}_{\nu}^{\alpha} = 0$$

Oeduction of the vacuum spacetime conservation laws :

$$\overline{
abla}_eta \overline{G}^lpha_eta = 0, \qquad \overline{G}^lpha_eta := \overline{\mathcal{R}}^lpha_eta - rac{\overline{\mathcal{R}}}{2} \; \delta^lpha_eta$$

4. Gravitation with sources : conservation laws

Variational formulation of Hilbert-Einstein action with source:

$$\delta\mathscr{S} := \frac{1}{2\chi} \int_{\mathscr{B}} \left(\overline{\mathfrak{R}}^{\alpha\beta} - \frac{\overline{\mathcal{R}}}{2} g^{\alpha\beta} \right) \mathcal{L}_{\xi} g_{\alpha\beta} \overline{\omega}_{n} + \int_{\mathscr{B}} \sigma^{\alpha\beta} \mathcal{L}_{\xi} g_{\alpha\beta} \overline{\omega}_{n} = 0$$

2 Field equation and **energy-momentum** tensor $\sigma^{\alpha\beta}$:

$$(1/2\chi)\left(\overline{\Re}^{\alpha\beta}-\overline{\mathcal{R}}/2\ g^{\alpha\beta}\right)+\sigma^{\alpha\beta}=0,$$

Solution laws following $\delta g_{\alpha\beta} \longrightarrow \mathcal{L}_{\xi} g_{\alpha\beta} = \overline{\nabla}_{\alpha} \xi_{\beta} + \overline{\nabla}_{\beta} \xi_{\alpha}$:

$$\overline{\nabla}_{\beta}\sigma^{\alpha\beta} + (1/2\chi)\overline{\nabla}_{\beta}\left(\overline{\mathfrak{R}}^{\alpha\beta} - \overline{\mathcal{R}}/2 \ g^{\alpha\beta}\right) = 0 \quad \Longrightarrow \quad \overline{\nabla}_{\beta}\sigma^{\alpha\beta} = 0$$

The conservation laws of σ "do not contain" the gravitational term.

(1) Two-steps approach similar to use **substantial variation** which includes transformations of space(time) coordinates and of physical fields (e.g. <u>Obukhov et al. 2014-5</u>, Lompay & Petrov, 2013, ...)

(2) Spacetime in Einstein RG is a second gradient continuum :

$$\begin{aligned} \overline{\Re}^{\lambda}_{\alpha\beta\mu} &= (1/2)g^{\lambda\sigma} \left(\partial_{\mu}\partial_{\alpha}g_{\sigma\beta} - \partial_{\mu}\partial_{\beta}g_{\sigma\alpha} + \partial_{\sigma}\partial_{\beta}g_{\mu\alpha} - \partial_{\sigma}\partial_{\alpha}g_{\mu\beta}\right) \\ &+ (1/4)g^{\lambda\sigma} \left(\partial_{\alpha}g_{\sigma\gamma} + \partial_{\gamma}g_{\alpha\sigma} - \partial_{\sigma}g_{\alpha\gamma}\right)g^{\gamma\kappa} \left(\partial_{\beta}g_{\kappa\mu} + \partial_{\mu}g_{\beta\kappa} - \partial_{\kappa}g_{\beta\mu}\right) \\ &- (1/4)g^{\lambda\sigma} \left(\partial_{\beta}g_{\sigma\gamma} + \partial_{\gamma}g_{\beta\sigma} - \partial_{\sigma}g_{\beta\gamma}\right)g^{\gamma\kappa} \left(\partial_{\alpha}g_{\kappa\mu} + \partial_{\mu}g_{\alpha\kappa} - \partial_{\kappa}g_{\alpha\mu}\right) \end{aligned}$$

Remarque

The goal is now to extend the conservation laws which takes the generic expression $\nabla_{\beta}\sigma^{\alpha\beta} = 0$ to more general geometries of **Relative** Gravitation and to some Inelastic continuum.

III. Description of some Inelastic Deformation

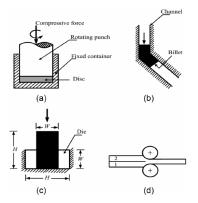


Figure: Material Processing with **Grain Refinement** by Severe Plastic Deformation : Strengthening of material.

Image: A math and A

Severe Plastic deformation : Example of HPT

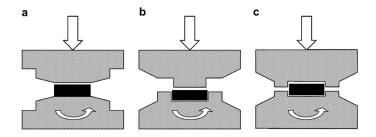
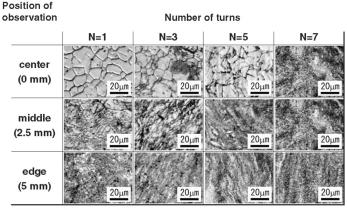


Figure: **Hight Pressure Torsion** Loadings : (a) Unconstrained; (b) and (c) Constrained.

Exemples of HPT plastification



(distance from center)

Figure: Optical Micrographs : Microstructures after HPT at the **center**, **half-way** position and **edge** of the disk in a magnesium AZ61 alloy after processing *N* turns at 423 K (Zhilayev & Langdon 2008).

Severe Plastic Deformation

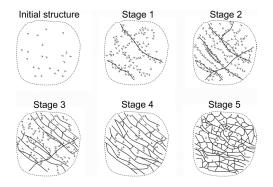


Figure: Grain size refinement under SPD (Cao et al. 2018):

(1) Initial : formation of large size dislocations cell blocks containing dislocations and dislocation cells structures;

(2) Formation of micro-bands and transformation of some nearly dislocations cells into cells blocks;

(3) Formation of lamellar sub-grains containing numerous dislocations;

(4) Formation of well-developed lamellar sub-grains and some equiaxed sub-grains

(5) Homogeneous distribution of equiaxed ultrafine grains or nano-grains..

Plastic deformation : Some conclusions

- **Twinning** and **relative motions** of grain with refinement (e.g. Zhylaiev & Langdon 2008) are mostly the mechanisms underlying plastic deformation.
- In addition, density of dislocation augments when the plastic deformation increases (e.g. Cao et al. 2018).
- Asympotic response. For complicated deformation operating in SPD, microstructure mostly reach a steady state at which further deformation does not change the overall microstructure.

Dislocations density and grains relative motions

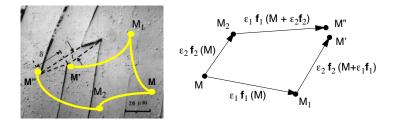


Figure: (L) Shear band formation. (R) Cartan parallelogram.

- (Left) Illustration of a shear band formation resulting from a **defect discontinuity** (δ, ℓ) at points M' and M''.
- (Right) Cartan parallelogram with area ε₁ × ε₂. At the continuum level, nucleation and migration of a great number of defects (discontinuities) occur, and give rise to macroscopic plastic strain (Lüders-Hartmann bands)

Dislocations density and grains relative motions

Schouten's relations, 1954:

Theorem

Say an affinely connected manifold \mathscr{B} . Let consider two arbitrary vectors \mathbf{f}_1 and \mathbf{f}_2 at a point M, they define two paths of length ϵ_1 and ϵ_2 . Then:

$$\begin{cases} \lim_{(\epsilon_1,\epsilon_2)\to 0} \frac{(\theta'-\theta'')}{\epsilon_1\epsilon_2} = \aleph(\mathbf{f}_1,\mathbf{f}_2)[\theta] \\ \lim_{(\epsilon_1,\epsilon_2)\to 0} \frac{(\mathbf{w}''-\mathbf{w}')}{\epsilon_1\epsilon_2} = \Re(\mathbf{f}_1,\mathbf{f}_2,\mathbf{w}) - \nabla_{\aleph(\mathbf{f}_1,\mathbf{f}_2)}\mathbf{w} \end{cases}$$

Proof (Partly R 1997, complete R 2021) With

$$\begin{split} & \aleph\left(\mathbf{f}_{1},\mathbf{f}_{2}\right)\left[\theta\right] \quad := \quad \left(\nabla_{\mathbf{f}_{1}}\mathbf{f}_{2}-\nabla_{\mathbf{f}_{2}}\mathbf{f}_{1}\right)\left[\theta\right]-\left[\mathbf{f}_{1},\mathbf{f}_{2}\right]\left[\theta\right] \\ & \Re\left(\mathbf{f}_{1},\mathbf{f}_{2},\mathbf{w}\right) \quad := \quad \nabla_{\mathbf{f}_{1}}\nabla_{\mathbf{f}_{2}}\mathbf{w}-\nabla_{\mathbf{f}_{2}}\nabla_{\mathbf{f}_{1}}\mathbf{w}-\nabla_{\left[\mathbf{f}_{1},\mathbf{f}_{2}\right]}\mathbf{w} \end{split}$$

Sketch of the proof (1)

1) Consider a scalar field $\theta(M)$, we have respectively the relations:

$$\begin{cases} \theta(M_1) = \theta(M + \varepsilon_1 \mathbf{f}_1) = \theta(M) + \nabla_{\varepsilon_1 \mathbf{f}_1(M)} \theta(M) \\ \theta(M') = \theta(M_1 + \varepsilon_2 \mathbf{f}_2(M + \varepsilon_1 \mathbf{f}_1)) = \theta(M_1) + \nabla_{\varepsilon_2 \mathbf{f}_2(M + \varepsilon_1 \mathbf{f}_1)} \theta(M_1) \end{cases}$$

We obtain the value of scalar field at M' in terms of **its value at** M, by noticing $\theta(M') = \theta'$ and $\theta(M) = \theta$, :

$$\begin{aligned} \theta' &= \theta + \varepsilon_1 \nabla_{\mathbf{f}_1} \theta + \varepsilon_2 \nabla_{\mathbf{f}_2} \theta + \varepsilon_2 \varepsilon_1 \nabla_{\nabla_{\mathbf{f}_1} \mathbf{f}_2} \theta + \varepsilon_2 \varepsilon_1 \nabla_{\mathbf{f}_2} \nabla_{\mathbf{f}_1} \theta \\ &+ \varepsilon_2 \varepsilon_1^2 \nabla_{\nabla_{\mathbf{f}_1} \mathbf{f}_2} \nabla_{\mathbf{f}_1} \theta \end{aligned}$$
 (1)

2) Similarly, we also obtain:

$$\theta'' = \theta + \varepsilon_2 \nabla_{\mathbf{f}_2} \theta + \varepsilon_1 \nabla_{\mathbf{f}_1} \theta + \varepsilon_1 \varepsilon_2 \nabla_{\nabla_{\mathbf{f}_2} \mathbf{f}_1} \theta + \varepsilon_1 \varepsilon_2 \nabla_{\mathbf{f}_1} \nabla_{\mathbf{f}_2} \theta + \varepsilon_1 \varepsilon_2^2 \nabla_{\nabla_{\mathbf{f}_2} \mathbf{f}_1} \nabla_{\mathbf{f}_2} \theta$$
(2)

Sketch of the proof (2)

3) Following the same method, say a vector field, with $\mathbf{w} := \mathbf{w}(M)$,:

$$\mathbf{w}' = \mathbf{w} + \varepsilon_1 \nabla_{\mathbf{f}_1} \mathbf{w} + \varepsilon_2 \nabla_{\mathbf{f}_2} \mathbf{w} + \varepsilon_2 \varepsilon_1 \nabla_{\nabla_{\mathbf{f}_1} \mathbf{f}_2} \mathbf{w} + \varepsilon_2 \varepsilon_1 \nabla_{\mathbf{f}_2} \nabla_{\mathbf{f}_2} \nabla_{\mathbf{f}_1} \mathbf{w}$$

$$+ \varepsilon_2 \varepsilon_1^2 \nabla_{\nabla_{\mathbf{f}_1} \mathbf{f}_2} \nabla_{\mathbf{f}_1} \mathbf{w}$$

$$(3)$$

and

$$\mathbf{w}'' = \mathbf{w} + \varepsilon_2 \nabla_{\mathbf{f}_2} \mathbf{w} + \varepsilon_1 \nabla_{\mathbf{f}_1} \mathbf{w} + \varepsilon_1 \varepsilon_2 \nabla_{\nabla_{\mathbf{f}_2} \mathbf{f}_1} \mathbf{w} + \varepsilon_1 \varepsilon_2 \nabla_{\mathbf{f}_1} \nabla_{\mathbf{f}_2} \mathbf{w}$$

$$+ \varepsilon_1 \varepsilon_2^2 \nabla_{\nabla_{\mathbf{f}_2} \mathbf{f}_1} \nabla_{\mathbf{f}_2} \mathbf{w}$$

$$(4)$$

4) Final step:

(1) - (2) induces the first relation on torsion.

(3) - (4) induces the second result on curvature (and torsion). \Box

Remarque

Calculus is done **exculsively** at point M (then $\mathcal{T}_M \mathcal{B}$). Without curvature but with torsion, we may have vector field discontinuities.

Covariance of Lagrangian \mathscr{L} (Antonio & R 2011)

Theorem

The model of continuum defined by the Lagrangian (Reichenbach, 1929)

 $\mathscr{L} = \mathscr{L}(g_{\alpha\beta}, \ \mathsf{\Gamma}^{\gamma}_{\alpha\beta}, \ \partial_{\lambda}\mathsf{\Gamma}^{\gamma}_{\alpha\beta})$

is **covariant** if and only if

$$\mathscr{L} = \mathscr{L}(\mathsf{g}_{\alpha\beta}, \ \aleph^{\gamma}_{\alpha\beta}, \ \mathfrak{R}^{\gamma}_{\alpha\beta\lambda})$$

Remarks :

- **Primal / internal variables** are metric $g_{\alpha\beta}$, torsion $\aleph_{\alpha\beta}^{\gamma}$, and curvature $\Re_{\alpha\beta\lambda}^{\gamma}$ (**Continuum physics** : elasticity, fluid mechanics, gravitation, electromagnetism, plastic deformation ...)
- This theorem extends **Cartan** (1922) and **Lovelock-Rund** (1971, 1975) theorems from Riemann to **Riemann-Cartan continuum**.

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Sketch of proof

Proof: Different steps is given below (Antonio & R, 2011):

(Consider an arbitrary change of coordinates (C^{∞} diffeomorphism)

$$x^{lpha} = x^{lpha} \left(y^{i}
ight), \quad ext{and} \quad J^{lpha}_{i} := \partial_{i} x^{lpha}, \cdots$$

Write as follows : $g_{ij} = J_i^{\alpha} J_j^{\beta} g_{\alpha\beta}$, ...

2 Assume diffeomorphism-invariance (covariance) and the transformation rules for the metric $g_{\alpha\beta}$ and the connection $\Gamma^{\gamma}_{\alpha\beta}$

$$\mathscr{L}(g_{ij}, \Gamma_{ij}^{k}, \partial_{l}\Gamma_{ij}^{k} + \Gamma_{ij}^{m}\Gamma_{lm}^{k}) = \mathscr{L}(g_{\alpha\beta}, \Gamma_{\alpha\beta}^{\gamma}, \partial_{\lambda}\Gamma_{\alpha\beta}^{\gamma} + \Gamma_{\alpha\beta}^{\mu}\Gamma_{\lambda\mu}^{\gamma})$$

- **3** Decompose $\Gamma^{\gamma}_{\alpha\beta}$ and $\partial_{\alpha}\Gamma^{\kappa}_{\beta\lambda} + \Gamma^{\xi}_{\beta\lambda}\Gamma^{\kappa}_{\alpha\xi}$ into symmetric and skew symmetric parts.
- **O Differentiate** \mathscr{L} with respect to J_i^{α} , ...
- **③** Apply the **Quotient theorem** (h tensor): $A^{\alpha\beta}h_{\alpha\beta} + B^{\alpha\beta\gamma}\nabla_{\gamma}h_{\alpha\beta}$ scalar ⇒ $A^{\alpha\beta}$, $B^{\alpha\beta\gamma}$ tensors □.

Remarque

 ${\mathscr L}$ satisfies the second axiom of Hilbert on the invariance of the action.

Differential geometry links RG and IC

• Elastic Continuum (deduced from Cauchy-Weyl theorem, 1850, 1939)²

$$\mathscr{L} = \mathscr{L}(g_{\alpha\beta}), \qquad g_{\alpha\beta} := \mathbf{g}(\mathbf{f}_{\alpha}, \mathbf{f}_{\beta})$$

where $\{\mathbf{f}_{\alpha}, \alpha = 1, n\}$ is an embedded base of \mathscr{B} (Pfaffian **F**!).

• Einstein Relative Gravitation

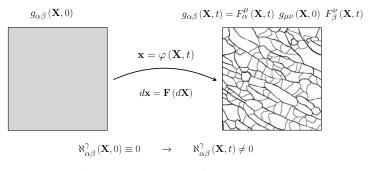
$$\mathscr{L} = \mathscr{L}(g_{\alpha\beta}, \overline{\mathfrak{R}}^{\gamma}_{\alpha\beta\lambda}) \quad \text{as} \quad \mathscr{L} = (1/2\chi) g^{\alpha\beta} \overline{\mathfrak{R}}^{\lambda}_{\lambda\alpha\beta}$$
extended to **Einstein-Cartan**
$$\mathscr{L} = \mathscr{L}\left(g_{\alpha\beta}, \aleph^{\gamma}_{\alpha\beta}, \Re^{\gamma}_{\alpha\beta\lambda}\right).$$

• Class of Inelastic Continua

$$\mathscr{L} = \mathscr{L}\left(\mathsf{g}_{lphaeta}, leph_{lphaeta}^{\gamma}, \Re_{lphaeta\lambda}^{\gamma}
ight)$$

²Complete proof in R2003 for rotational transformations @→ < ≡→ < ≡→ = ∽ < <

Some Inelastic continuum models $\mathscr{L} = \mathscr{L}(g_{\alpha\beta}, \aleph^{\gamma}_{\alpha\beta}, \Re^{\lambda}_{\alpha\beta\gamma})$



 $\Re^{\gamma}_{\alpha\beta\lambda}\left(\mathbf{X},0\right)\equiv0\qquad\rightarrow\qquad\Re^{\gamma}_{\alpha\beta\lambda}\left(\mathbf{X},t\right)\neq0$

Figure: Inelastic deformation includes : (1) variation of the metric components; (2) variations of the **torsion** and **curvature**.

IV. Principle of General Covariance and Applications

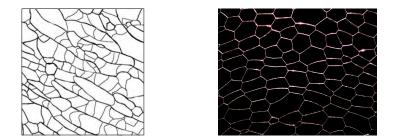


Figure: Left: Plastic deformation of (Metallic alloy, set of microcosms); Middle: Spacetime in Loop Quantum Gravitation (set of "quanta"). For the general case, we have.

Principle of General Covariance.³ (e.g. Souriau 1975, Duval & Künzle 1978, R 2018) The principle dwells upon Poincaré's invariant integral involving Lie derivatives,:

$$\delta\mathscr{S} = \int_{\mathscr{B}} \left(\left(\sigma^{\alpha\beta} \mathcal{L}_{\xi} g_{\alpha\beta} + \sum_{\gamma}^{\alpha\beta} \mathcal{L}_{\xi} \aleph^{\gamma}_{\alpha\beta} + \Xi^{\alpha\beta\mu}_{\lambda} \mathcal{L}_{\xi} \Re^{\lambda}_{\alpha\beta\mu} \right) \omega_{n} = 0$$

 ∇ compatible with ω_n and ξ divergence-free.

Conservation laws are obtained when the trajectory is shifted while the action is left unchanged. Active diffeomorphisms are defined by Lie derivative variations

$$\{\mathcal{L}_{\xi}g_{\alpha\beta}, \quad \mathcal{L}_{\xi}\aleph_{\alpha\beta}^{\gamma}, \quad \mathcal{L}_{\xi}\Re_{\alpha\beta\mu}^{\lambda}\}$$

³It can be extended to Noether-Klein method to provide Klein identities by considering ξ , $\nabla \xi$, $\nabla \nabla \xi$, ...)e.g. Lompay & Petrov 2013) $\in \mathbb{P} \times \mathbb{P} \times \mathbb{P} \times \mathbb{P} \times \mathbb{P}$

Application (1) : Lagrangian of the form $\mathscr{L}(g_{lphaeta})$

Constitutive laws. The "stress" σ is defined as :

$$\sigma^{\alpha\beta} := \frac{\partial \mathscr{L}}{\partial g_{\alpha\beta}} (g_{\mu\nu}, \underbrace{\aleph_{\mu\nu}^{\gamma}, \Re_{\mu\nu\epsilon}^{\gamma}}_{\text{frozen}})$$

Theorem

Consider a Riemann-Cartan continuum $(\mathcal{B}, \mathbf{g}, \nabla)$ with a Lagrangian function $\mathcal{L}(g_{\alpha\beta})$. Then :

$$\nabla_{\alpha}\sigma_{\gamma}^{\alpha} = \mathbf{0}$$

Proof: 1. Local invariance. The PGC of the Lagrangian holds:

$$\begin{split} \delta \mathscr{L} &:= \int \sigma^{\alpha\beta} \ \mathcal{L}_{\xi} g_{\alpha\beta} \ \omega_{n} \\ \mathcal{L}_{\xi} \ g_{\alpha\beta} &= \xi^{\gamma} \ \nabla_{\gamma} g_{\alpha\beta} + g_{\gamma\beta} \ \nabla_{\alpha} \xi^{\gamma} + g_{\alpha\gamma} \ \nabla_{\beta} \xi^{\gamma} \\ &+ \xi^{\gamma} \left(g_{\alpha\nu} \ \aleph^{\nu}_{\gamma\beta} + g_{\nu\beta} \ \aleph^{\nu}_{\gamma\alpha} \right) \end{split}$$

Integrate by parts and shift the boundary terms to obtain:

$$\nabla_{\alpha}\sigma_{\gamma}^{\alpha} + \aleph_{\alpha\gamma}^{\lambda} \ \sigma_{\lambda}^{\alpha} = \mathbf{0}$$

Proof (2)

2. Global Invariance. By considering a uniform vector field ξ such that $\nabla \xi \equiv 0$, we obtain a **Klein Identity**:

$$\sigma^{\alpha\beta}\left(g_{\alpha\nu} \aleph^{\nu}_{\gamma\beta} + g_{\nu\beta} \aleph^{\nu}_{\gamma\alpha}\right) = 2 \sigma^{\alpha}_{\mu} \aleph^{\mu}_{\gamma\alpha} \equiv 0$$

We deduce the classical continuum conservation laws:

$$abla_{\alpha}\sigma_{\gamma}^{\alpha} = \mathbf{0}$$
 \Box

- The implicit presence of the **torsion** within the covariant derivative should be accounted for (e.g. Futhazar et al. 2014). (Spatial fading of waves)
- Other identities might exist according to the choice of ξ (x). (e.g. Lompay & Petrov 2013)

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Example of 3D-elasticity with non evolving defects

 $\textcircled{0} \ \mathsf{Lagrangian} \ \mathscr{L}$

$$u(\varepsilon) = \frac{\lambda}{2} \operatorname{Tr}^2 \varepsilon + \mu \operatorname{Tr}(\varepsilon^2), \qquad \mathscr{L} := \frac{\rho}{2} \| \frac{d\mathbf{v}}{dt} \|^2 - u(\varepsilon)$$

2 Small strain assumption:

$$g_{lphaeta} \simeq \delta_{lphaeta} +
abla_{lpha} u_{eta} +
abla_{eta} u_{lpha}, \qquad arepsilon := rac{1}{2} \left(
abla \mathbf{u} +
abla \mathbf{u}^{\mathrm{T}}
ight)$$

3 Extended Navier equation:

$$\partial_t^2 u^{\alpha} = \frac{\lambda + \mu}{\rho} \nabla^{\alpha} \nabla_{\beta} u^{\beta} + \frac{\mu}{\rho} \nabla^{\beta} \nabla_{\beta} u^{\alpha} + \frac{\mu}{\rho} \hat{g}^{\alpha\mu} \left(\Re_{\mu\nu} u^{\gamma} - \aleph_{\beta\mu}^{\gamma} \nabla_{\gamma} u^{\beta} \right)$$

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Remarque

Wave behaviour similar to the bending of light (electromagnetic wave) waving through a gravitational field !

Application (2) : Inelastic continuum

1 Separate the torsion into skew-symmetric **Cartan's forms**:

$$\aleph^{\gamma}_{\alpha\beta} \quad \rightarrow \quad \left\{ \omega^{(1)}_{\alpha\beta}, \omega^{(2)}_{\alpha\beta}, \omega^{(3)}_{\alpha\beta} \right\}$$

(dislocations density)

2 Reduction of curvature to **Ricci tensor** in 3*D*:

$$\Re^{\gamma}_{\lambdalphaeta} \quad o \quad \mathcal{R}_{lphaeta} := \Re^{\gamma}_{\gammalphaeta}$$

(disclinations density)

(Covariant) Lagrangian of some class of inelastic continua

$$\mathscr{L} = \mathscr{L}(g_{\alpha\beta}, \omega_{\alpha\beta}^{(n)}, \mathcal{R}_{\alpha\beta})$$

1 Gauge Invariance (divergence-free gauge ξ)

$$\delta\mathscr{S} = \int_{\mathscr{B}} \left(\sigma^{\alpha\beta} \mathcal{L}_{\xi} g_{\alpha\beta} + \sum_{(n)}^{\alpha\beta} \mathcal{L}_{\xi} \omega_{\alpha\beta}^{(n)} + \Xi^{\alpha\beta} \mathcal{L}_{\xi} \mathcal{R}_{\alpha\beta} \right) \omega_n = 0$$

2 Constitutive laws

$$\sigma^{\alpha\beta} := \frac{\partial \mathscr{L}}{\partial g_{\alpha\beta}}, \qquad \Sigma^{\alpha\beta}_{(n)} := \frac{\partial \mathscr{L}}{\partial \omega^{(n)}_{\alpha\beta}}, \qquad \Xi^{\alpha\beta} := \frac{\partial \mathscr{L}}{\partial \mathcal{R}_{\alpha\beta}}$$

Remarque

Stress $\sigma^{\alpha\beta}$ is symmetric, $\sum_{(n)}^{\alpha\beta}$ are skew-symmetric. The dual of Ricci curvature $\Xi^{\alpha\beta}$ is not necessarily symmetric.

Conservation Laws from Gauge Theory

Theorem

Let \mathscr{B} a continuum with Lagrangian $\mathscr{L}(g_{\alpha\beta}, \omega_{\alpha\beta}^{(n)}, \mathcal{R}_{\alpha\beta}), n = 1, 2, 3$. Then the conservation laws hold:

$$\nabla_{\alpha}\tilde{\sigma}^{\alpha}_{\gamma}=\mathbf{0}$$

where we define the generalized stress $\tilde{\sigma}^{\alpha}_{\gamma}$:

$$2\tilde{\sigma}_{\gamma}^{\alpha} := 2 \,\, \sigma_{\gamma}^{\alpha} + 2 \, \sum_{(n)}^{\alpha\beta} \omega_{\gamma\beta}^{(n)} + 2 \, \Xi^{\alpha\beta} \mathcal{R}_{\gamma\beta}$$

Remarque

Presence of dislocations density in $\tilde{\sigma}^{\alpha}_{\gamma}$ seems giving new interesting insights for modeling strain hardening phenomenon during plastic deformation : "the first problem to be attempted by dislocation theory and may be the last to be solved" (e.g. Kocks & Mecking, 2003).

Sketch of the proof

Lie derivatives of primal variables:

$$\begin{split} \mathcal{L}_{\xi} \mathbf{g}_{\alpha\beta} &= \xi^{\gamma} \left(\nabla_{\gamma} \mathbf{g}_{\alpha\beta} + \mathbf{g}_{\alpha\nu} \, \aleph_{\gamma\beta}^{\nu} + \mathbf{g}_{\nu\beta} \, \aleph_{\gamma\alpha}^{\nu} \right) + \left(\mathbf{g}_{\gamma\beta} \, \nabla_{\alpha} \xi^{\gamma} + \mathbf{g}_{\alpha\gamma} \, \nabla_{\beta} \xi^{\gamma} \right) \\ \mathcal{L}_{\xi} \omega_{\alpha\beta}^{(n)} &= \xi^{\gamma} \left(\nabla_{\gamma} \omega_{\alpha\beta}^{(n)} + \omega_{\nu\beta}^{(n)} \, \aleph_{\gamma\alpha}^{\nu} + \omega_{\alpha\nu}^{(n)} \, \aleph_{\gamma\beta}^{\nu} \right) + \left(\omega_{\gamma\beta}^{(n)} \nabla_{\alpha} \xi^{\gamma} + \omega_{\alpha\gamma}^{(n)} \, \nabla_{\beta} \xi^{\gamma} \right) \\ \mathcal{L}_{\xi} \mathcal{R}_{\alpha\beta} &= \xi^{\gamma} \left(\nabla_{\gamma} \mathcal{R}_{\alpha\beta} + \mathcal{R}_{\nu\beta} \, \aleph_{\gamma\alpha}^{\nu} + \mathcal{R}_{\alpha\nu} \, \aleph_{\gamma\beta}^{\nu} \right) + \left(\mathcal{R}_{\gamma\beta} \nabla_{\alpha} \xi^{\gamma} + \mathcal{R}_{\alpha\gamma} \nabla_{\beta} \xi^{\gamma} \right) \end{split}$$

2 The three Klein-Noether identities take the following form for $\gamma = 1, 2, 3$:

$$\begin{array}{lll} 0 & = & \sigma^{\alpha\beta} \left(\nabla_{\gamma} g_{\alpha\beta} + g_{\alpha\nu} \aleph^{\nu}_{\gamma\beta} + g_{\nu\beta} \aleph^{\nu}_{\gamma\alpha} \right) \\ & + & \sum_{n=1}^{n=3} \left[\Sigma^{\alpha\beta}_{(n)} \left(\nabla_{\gamma} \omega^{(n)}_{\alpha\beta} + \omega^{(n)}_{\nu\beta} \aleph^{\nu}_{\gamma\alpha} + \omega^{(n)}_{\alpha\nu} \aleph^{\nu}_{\gamma\beta} \right) \right. \\ & + & \Xi^{\alpha\beta} \left(\nabla_{\gamma} \mathcal{R}_{\alpha\beta} + \mathcal{R}_{\nu\beta} \aleph^{\nu}_{\gamma\alpha} + \mathcal{R}_{\alpha\nu} \aleph^{\nu}_{\gamma\beta} \right) \end{array}$$

By accounting fot these identities, it is easy to deduce the following equation by integrating by parts:

$$\int_{\mathscr{B}} \xi^{\gamma} \nabla_{\alpha} \left[2 \sigma_{\gamma}^{\alpha} + 2 \Sigma_{(n)}^{\alpha\beta} \omega_{\gamma\beta}^{(n)} + (\Xi^{\alpha\beta} \mathcal{R}_{\gamma\beta} + \Xi^{\beta\alpha} \mathcal{R}_{\beta\gamma}) \right] \omega_{n} = 0 \qquad \Box$$

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Comment on Generalized Stress

Remarque

Different names of $\sum_{(n)}^{\alpha\beta}$ and $\Xi^{\alpha\beta}$: Hypermomentum in Relative Gravitation with spin (e.g. Hehl et al., 1975-76,77,78), or Total energy stress (e.g. Duval and Künzle, 1978); Micro-stress and Polar micro-stress in Strain Gradient Plasticity (e.g. Gurtin and Anand, 2004-2005); Currents in Generalized Gravity (e.g. Obukhov et al. 2013, 2015, Lompay et al. 2013)

Remarque

Generalized stress $2\tilde{\sigma}^{\alpha}_{\gamma}$ includes

- **1** a "reversible" part σ_{γ}^{α}
- **2** contribution of **Cartan** forms $\omega_{\alpha\beta}^{(n)}$ (corresponding to a spin, due to dislocations) and its dual $\sum_{\alpha\beta}^{\alpha\beta}$
- **3** contribution of **Ricci** tensor $\mathcal{R}_{\alpha\beta}$ and its dual $\Xi^{\alpha\beta}$ (due to disclinations).

V. Concluding remarks



Figure: Plasticity, Gravitation and way of Life

The **Principle of General Covariance** (Gauge Theory) can be applied for Relative Gravitation spacetime and for Inelastic Continuum modeled as Riemann-Cartan manifold. Application in sequence the two invariance criteria is worth:

- **Covariance** (passive diffeomorphism, necessary criterium for any physics theory, e.g. Norton 2005)
- Principle of General Covariance (active gauge invariance, e.g. Souriau, 1975)

Some other applications (R 2023)

The same **sequential method** was applied to other various model :

- **Q** Riemann-Cartan spacetime / continuum $\mathscr{L}(g_{\alpha\beta}, \aleph^{\gamma}_{\alpha\beta}, \Re^{\gamma}_{\alpha\beta\lambda})$ (MAG without nonmetricity) (e.g. Obukhov et al, 2015);
- **2** Weitzeböck Continuum with Lagrangian $\mathscr{L}(g_{\alpha\beta}, \mathcal{T}^{\gamma}_{\alpha\beta})$ where $\mathcal{T}^{\gamma}_{\alpha\beta} := \Gamma^{\gamma}_{\alpha\beta} \overline{\Gamma}^{\gamma}_{\alpha\beta}$ is the contortion tensor (similar to plasticity);
- **3** "Gradient Weitzeböck" Continuum with Lagrangian $\mathscr{L}(g_{\alpha\beta}, \mathcal{T}^{\gamma}_{\alpha\beta}, \overline{\nabla}\mathcal{T}^{\gamma}_{\alpha\beta})$ which is very similar to the Strain Gradient Plasticity (e.g. review Voyiadjis & Song, 2019, Gurtin-Anand, 2005)

We get an **unified framework** for conservation laws :

$$\nabla_{\alpha} \tilde{\sigma}^{\alpha}_{\gamma} = \mathbf{0} \qquad \qquad \tilde{\sigma} = \sigma + \mathscr{C}(\aleph, \Re, \Sigma, \Xi) \qquad \text{or} \qquad \tilde{\sigma} = \sigma + \mathscr{C}(\mathcal{T}, \overline{\nabla} \mathcal{T}, \Sigma, \Xi)$$

Very similar generalized stress was obtained in e.g. **Medina et al, 2019**, for studying Einstein-Cartan cosmologies, avoiding singularity.

Merci pour votre attention !

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